

Remarks on Exponential Stability of Solutions for the Compressible p -th Power Newtonian Fluid with Large Initial Data

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Abstract. In this paper, we establish the exponential stability of the global spherically and cylindrically symmetric solutions in H^i ($i = 1, 2, 4$) for the p -th power Newtonian fluid in multi-dimension with large initial data. The key point is that the smallness of initial data is not needed if the initial data are cylindrically symmetric.

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1 Introduction

In this paper, we establish the exponential stability of global solutions in H^i ($i = 1, 2, 4$) for the compressible Navier-Stokes equations which describe the motion of the p -th power Newtonian fluid. We assume that pressure P , in terms of the density ρ and absolute temperature θ , is given by $P = R\rho^p\theta$ with constant $R > 0$ and the pressure exponent $p \geq 1$, and the specific internal energy $e = c_v\theta$ with constant specific heat coefficient $c_v > 0$. We assume that the corresponding solutions only depend on the radial variable r and the time variable $t \in [0, +\infty)$. Here we use $\vec{U}(\mathbf{x}, t)$ to denote the velocity of fluid and $\Omega =$

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$\{\mathbf{x}|0 < a \leq |\mathbf{x}| \leq b < +\infty\}$ is a symmetric domain in \mathbb{R}^N . As in [1–3], the equations can be read as

$$\begin{cases} \rho_t + (\rho u)_r + \frac{m\rho u}{r} = 0, \\ \rho \left(u_t + uu_r - \frac{v^2}{r} \right) + P_r = \beta \left(u_r + \frac{mu}{r} \right)_r, \\ \rho \left(v_t + uv_r + \frac{uv}{r} \right) = \mu \left(v_r + \frac{mv}{r} \right)_r, \\ \rho (w_t + uw_r) = \mu \left(w_{rr} + \frac{mw_r}{r} \right), \\ c_v \rho (\theta_t + u\theta_r) + P \left(u_r + \frac{mu}{r} \right) = k \left(\theta_{rr} + \frac{m\theta_r}{r} \right) + Q, \end{cases} \quad (1.1)$$

where the viscosity coefficients μ, λ satisfy the natural restrictions $\mu > 0, N\lambda + 2\mu \geq 0$ and the coefficient of heat conduction k is positive $k > 0, \beta = 2\mu + \lambda$ and

$$Q = \lambda \left(u_r + \frac{mu}{r} \right)^2 + \mu \left\{ \left(v_r - \frac{mv}{r} \right)^2 + w_r^2 + 2u_r^2 + \frac{2mu^2}{r^2} \right\}.$$

In the spherically symmetric case, $m = N - 1, r = |\mathbf{x}|, \vec{\mathbf{U}}(\mathbf{x}, t) = u(r, t)\frac{\mathbf{x}}{r}, v = w = 0$. In the cylindrically symmetric case, $m = 1, r = \sqrt{x_1^2 + x_2^2}$, and

$$\vec{\mathbf{U}}(\mathbf{x}, t) = u(r, t)\frac{(x_1, x_2, 0)}{r} + v(r, t)\frac{(-x_2, x_1, 0)}{r} + w(r, t)(0, 0, 1).$$

The following boundary and initial conditions can be given by

$$(u, v, w, \theta_r)(a, t) = (u, v, w, \theta_r)(b, t) = 0, \quad t \in [0, +\infty), \quad (1.2)$$

$$(\rho, u, v, w, \theta)(r, 0) = (\rho_0, u_0, v_0, w_0, \theta_0)(r), \quad r \in [a, b]. \quad (1.3)$$

As in [1, 4, 5], we denote by $\eta = \frac{1}{\rho}$ the specific volume of the flows and can transfer the problem (1.1)-(1.3) into the equations in Lagrangian coordinates as follows

$$\begin{cases} \eta_t = (r^m u)_x, & 0 < x < L, t > 0, \end{cases} \quad (1.4)$$

$$\begin{cases} u_t = r^m \left(-\frac{R\theta}{\eta^p} + \beta \frac{(r^m u)_x}{\eta} \right)_x + \frac{v^2}{r}, \end{cases} \quad (1.5)$$

$$\begin{cases} v_t = \mu r^m \left(\frac{(r^m v)_x}{\eta} \right)_x - \frac{uv}{r}, \end{cases} \quad (1.6)$$

$$\begin{cases} w_t = \mu r^m \left(\frac{(r^m w)_x}{\eta} \right)_x + \frac{\mu m \eta w}{r^2}, \end{cases} \quad (1.7)$$

$$\begin{cases} c_v \theta_t = \left(-\frac{R\theta}{\eta^p} + \beta \frac{(r^m u)_x}{\eta} \right) (r^m u)_x + \left(k \frac{r^{2m} \theta_x}{\eta} \right)_x + \mu \frac{(r^m v)_x^2}{\eta} \\ \quad + \mu \frac{r^{2m} w_x^2}{\eta} - 2\mu m (r^{m-1} u^2 + r^{m-1} v^2)_x. \end{cases} \quad (1.8)$$